

International Centre for Radio Astronomy Research RFI Assessment of Photovoltaic Modules for Radio Astronomy Applications

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THE UNIVERSITY OF WESTERN AUSTRALIA





- Introduction
- Interference model
- Preliminary measurements
- Conclusion



#### Introduction



#### Square Kilometre Array (SKA)

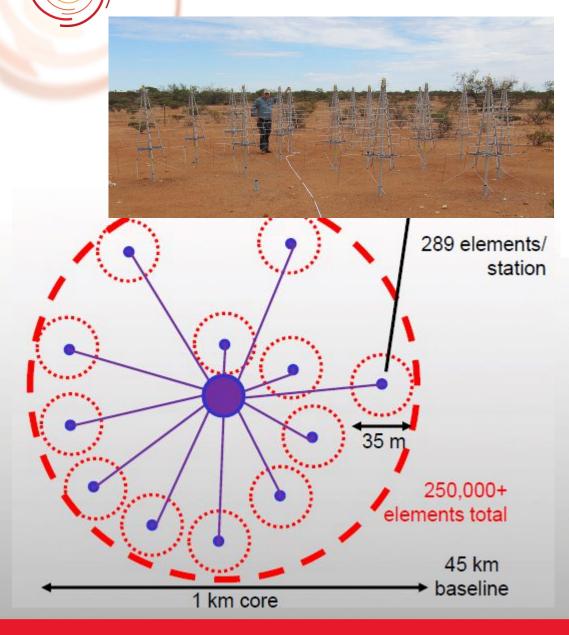
•Most sensitive radio telescope in the world

- International project
- •Collecting area 1 km<sup>2</sup>
- •Location: Australia-NZ, Southern Africa

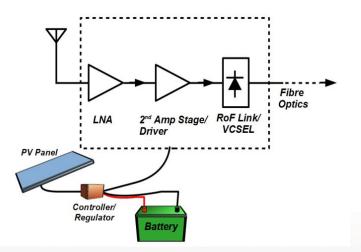
International Centre for Radio Astronomy Research (ICRAR) •The University of Western Australia + Curtin University •State (WA) and Federal support •Contributor to SKA-Low (AAVP)

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#### Introduction: SKA Low



CRA



# Each antenna element needs power for

- •LNA
- •2<sup>nd</sup> stage amplifier
- •RoF link

# Distributed photovoltaic modules:

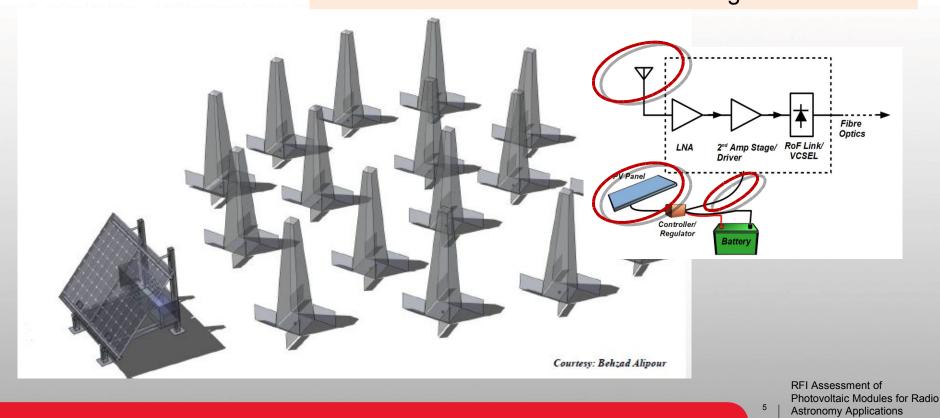
- •Solar panel
- Controller/Regulator
- Battery

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## Interference model

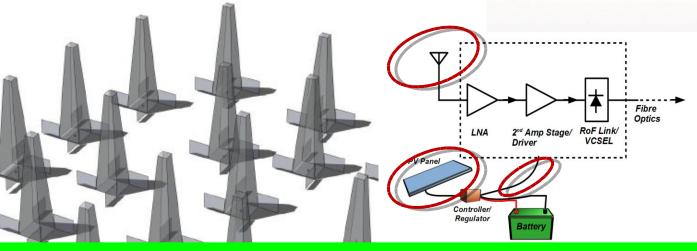
Controller/Regulator is the main noise source Conducted noise can interfere with performance of LNA Conducted noise can lead to radiation from: •Solar panel and the connecting cable •Battery and the connecting cable •SKA-Low antennas and the connecting cables





## Interference model

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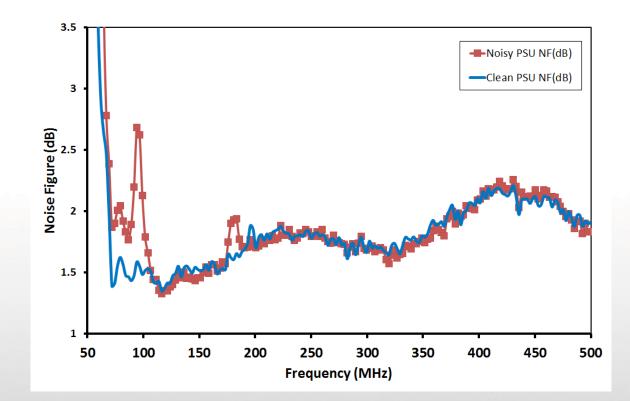


A compliance test of the complete system is not possible in a controlled lab environment. Therefore an assessment must be in form of a comprehensive analysis. **Tests on a component level don't give answers, but provide input for this analysis!** 

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#### Internal EMC



Increase of noise figure due to a noisy power supply

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#### Commercial PV modules



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http://store.solio.com/Solio-Store/CLASSIC2-Solar-Battery-Charger-S13-AF1RW

## **Commercial PV modules**

Solio - 100 kHz resolution bandwidth 40 Light OFF Light ON 30 Light ON - No load Mil-Std 461 Limit Field strength in dBuV/m Mil-Std -20 dB 20 -10 -20 80 40 60 100 120 140 160 180 200 Frequency in MHz

No difference between Light OFF and Light ON (charging) Connecting the load makes a difference!

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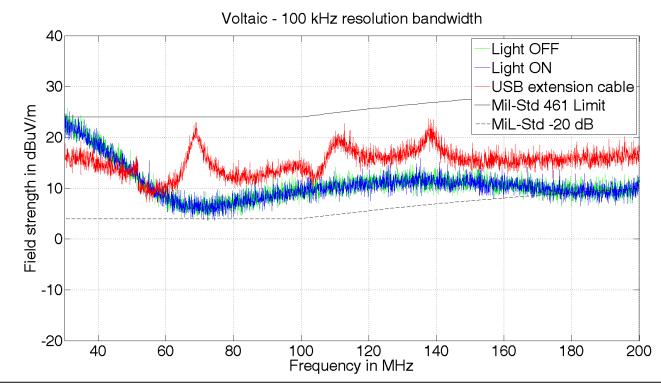
#### Commercial PV modules



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#### **Commercial PV modules**



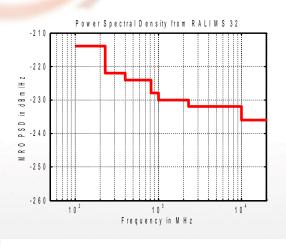
No difference between Light OFF and Light ON (charging) Connecting the load via an 0.5m long extension cable makes a difference!



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## **Selecting Limit Values**



Power spectral density (PSD) permitted at location of radio telescope antenna

#### 80 1L - S 461 D 70 IL-STD - 20dB (dBuV/m Ø. 10 km ) 60 d B u V /m MRO (dBuV/m Q 1 k m 50 .**\_** g th 4 stren 30 F ie Id **Emission limit values** 20 based on PSD values 10 for D > 1 km $10^{2}$ $10^{3}$ $10^{4}$ Frequency in MHz

Comparison of limit values

#### **MRO Requirements**

Distance > 10 km

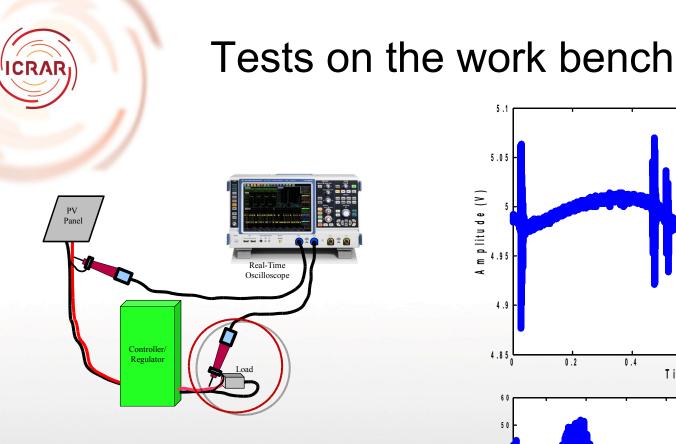
10 km > D > 1 km

1 km > D

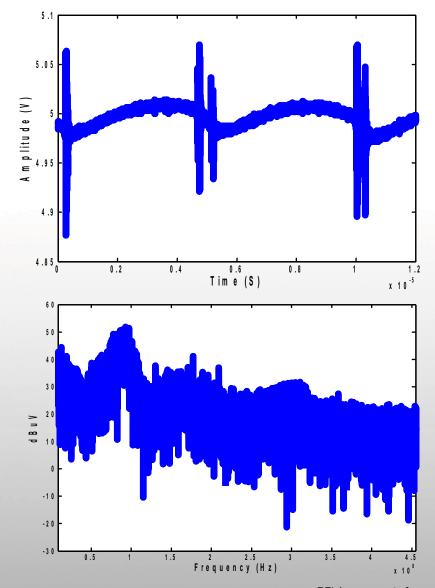
MIL-STD 461 Limits

- 1 km 20 dB below MIL-STD 461 Limits
  - 80 dB below MIL-STD 461 Limits (permitted only on a case by case basis)

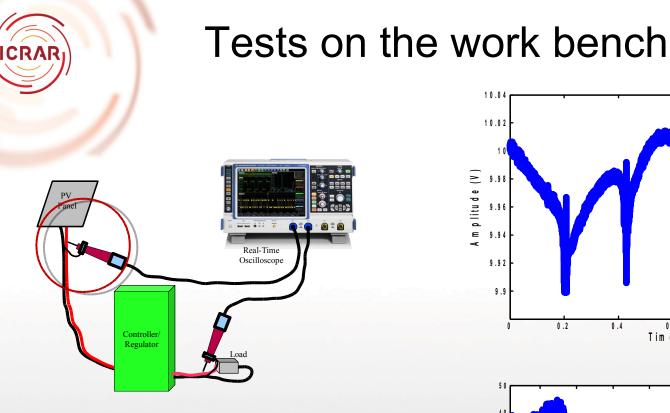
For SKA–Low there may be different limit values for different frequency ranges.

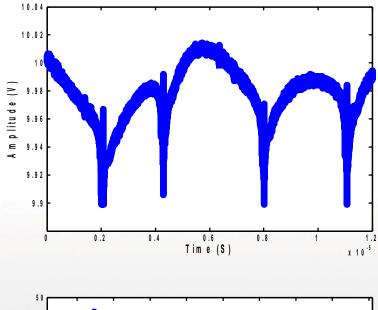


- RFI from a noisy SMPS: Ripple (low freq) and switching noise (high freq)
- Noise, including resonance effects, in the SKA-low band
- Reduce noise with filtering and selection of components
- Customize switch-mode regulator for SKA

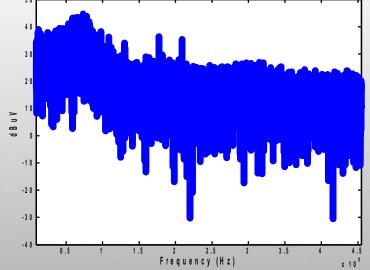


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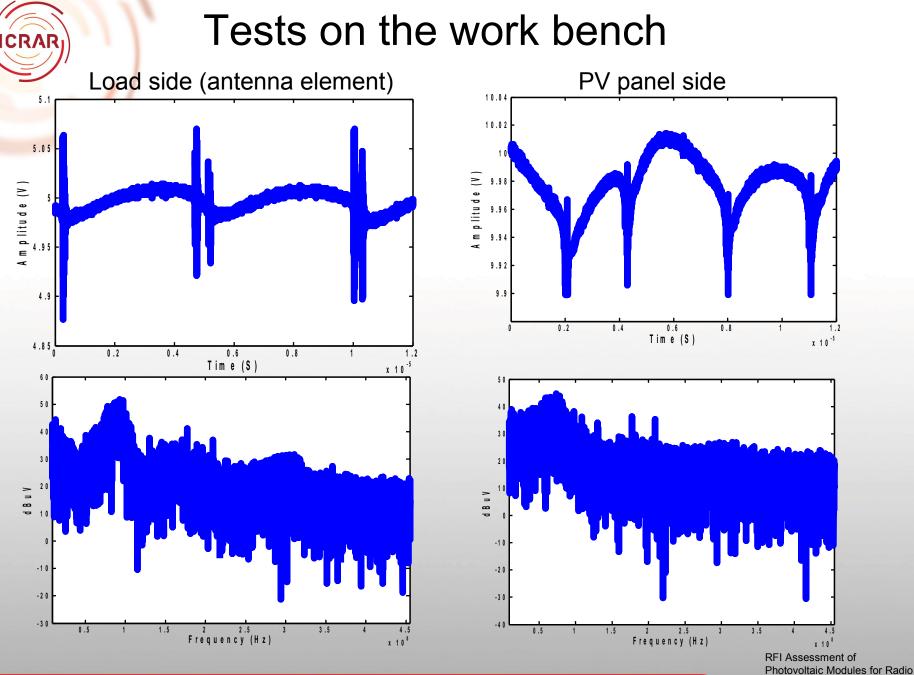




- Noise from the same SMPS •
- Noise, including resonance effects, in the ٠ SKA-low band
- Common to both sides: 175 MHz narrow ٠ band signal (possible ambient signal)
- Broadband noise is different at PV panel ٠ and load



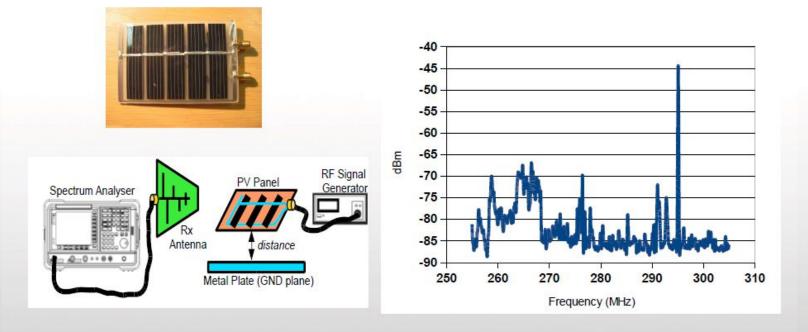
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Astronomy Applications



#### Tests on the work bench



- PV panel → Could potentially radiate EMI emission!
- · Metallic objects nearby interact with the panel
- · Grounding (distance and conductivity) has some influence

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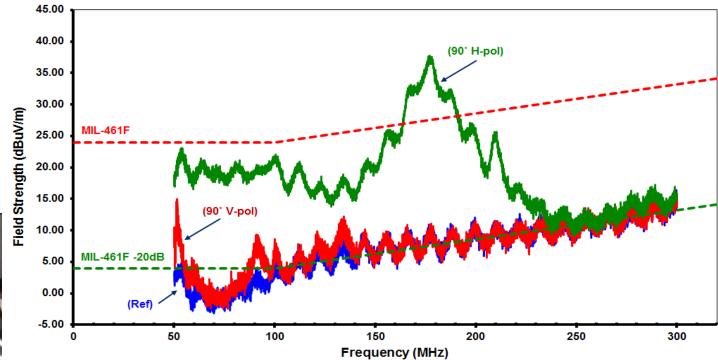
#### Tests in the EMC lab



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## Tests in the EMC lab





Horizontal polarisation much higher than vertical polarisation: Radiation from dipole antennas and/or connecting cables likely

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#### Conclusion

- Advantage of PV Solar modules
- Galvanic Isolation (with RoF links)
- Suitable for both SKA sites
- No need for power distribution network
- Preliminary EMI tests
- Emission from commercial-off-the-shelf equipment too high, custom design required
- Interaction from metallic objects near PV panel, and effect of long supply cables must be considered
- Tests on the work bench more convenient than measurements in EMC lab
- Preliminary tests on the work bench for comparison
- Determine Interface parameters (S11, conducted noise)

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#### Conclusion

- Prediction of RFI/EMI Strategy
- Determine antenna characteristics for: solar panels, SKA-low antennas, cabling (Analytically, simulation, measurements)
- Determine excitation based on: work bench measurements, EMC lab measurements and simulations
- Combine antenna characteristics and excitation
- Validate with measurements on representative configurations
- Purpose of EMC tests, in the early phase of the design
- Understanding the problem
- validation of the interference model
- Compliance assessment comes much later!

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#### Conclusion

- Purpose of a standard
- Conventional standard: in case of compliance, the likelihood of interference in a 'standard' environment is below a certain threshold
- SKA environment: a device, when integrated in the large system 'radio telescope' does not:
  - compromise the the performance of the telescope
  - compromise the performance of other telescopes on the same site
- In conventional standards much emphasis is placed on repeatability and reproducibility.
- In the SKA environment test results are more likely basis for a wider assessment, not just a yes/now compliance decision.

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